Introduction to IRIS data analysis

Tiago M. D. Pereira

Monday, April 11, 2016

14:00-15:30 Introduction to IRIS

15:30-16:30 Posters & Coffee

16:30–17:00 Solving questions

17:00-18:00 Hands-on tutorials

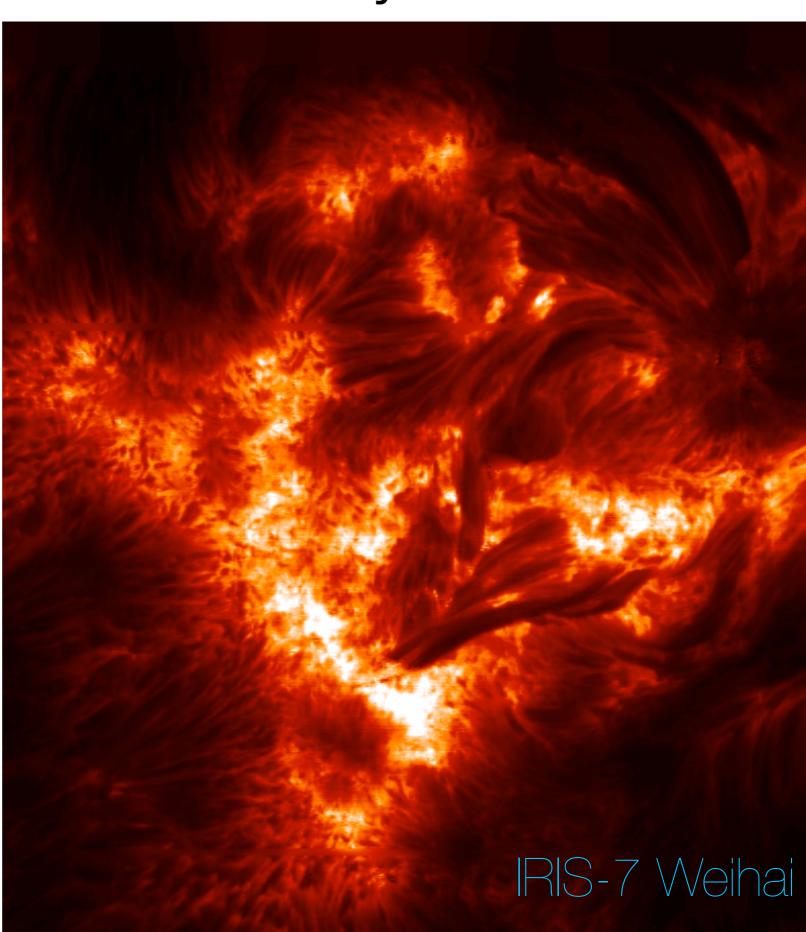
Tuesday, April 12, 2016

09:00-10:30 Hands-on tutorials









Lecture resources

Slides, notes, exercises:

http://folk.uio.no/tiago/iris7

lecture overview

Part 1

- Overview of IRIS, capabilities and resources
- Getting the data, quicklook tools
- Working with IRIS data

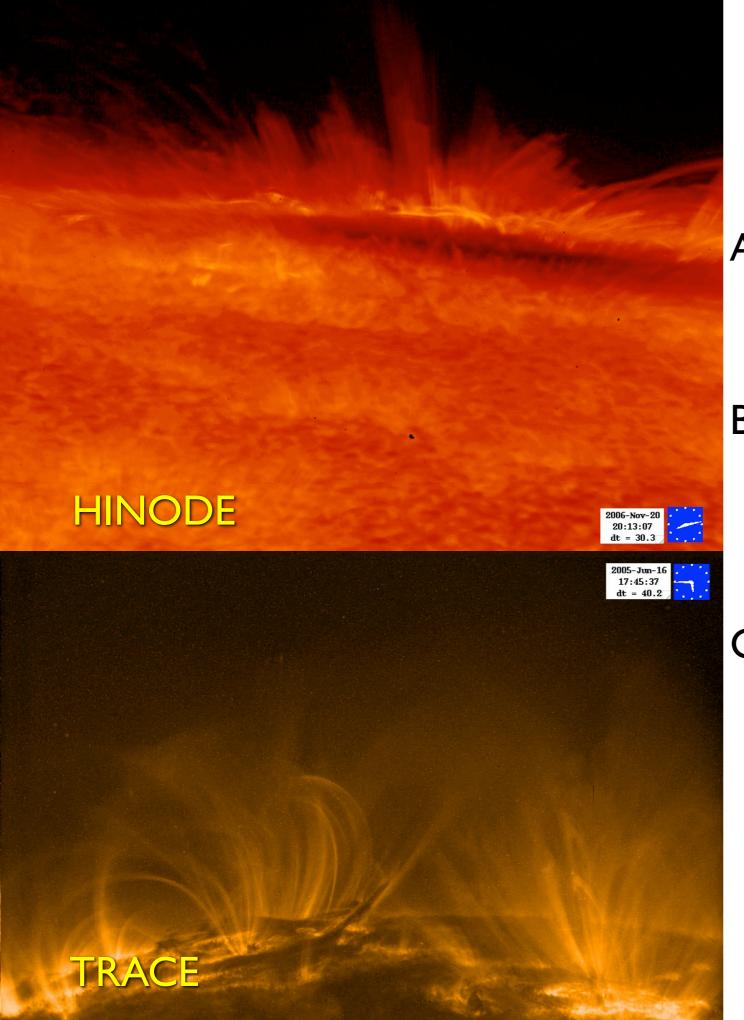
Part 2

- Additional Data Calibration
- Utility functions for Mg II lines
- CRISPEX
- Time to work on questions

Exercise questionsHands-on tutorials

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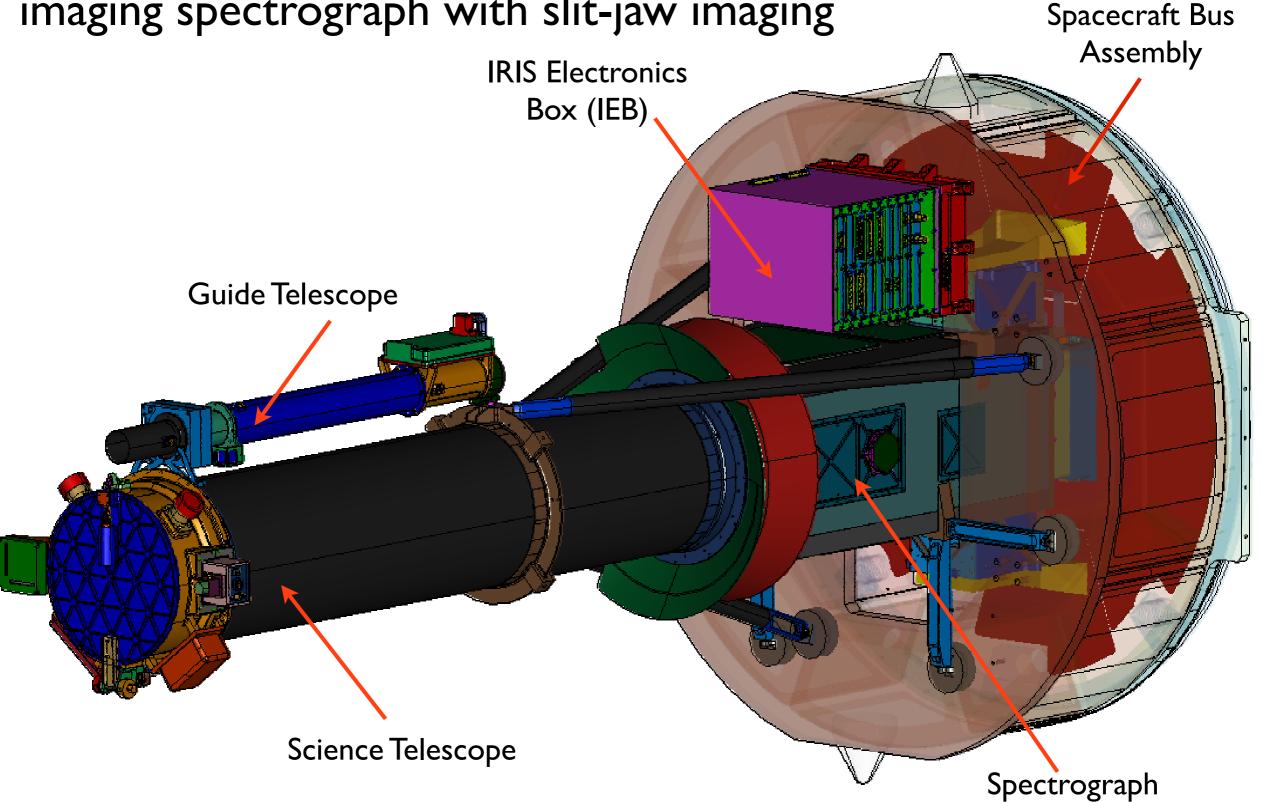
Major Science Goals of IRIS

- A. Which types of non-thermal energy dominate in the chromosphere and beyond?
- B. How does the chromosphere regulate the mass and energy supplied to the corona and heliosphere?
- C. How does magnetic flux and matter rise through the lower atmosphere, and what is the role of flux emergence in powering flares and mass ejections?

Courtesy Bart De Pontieu

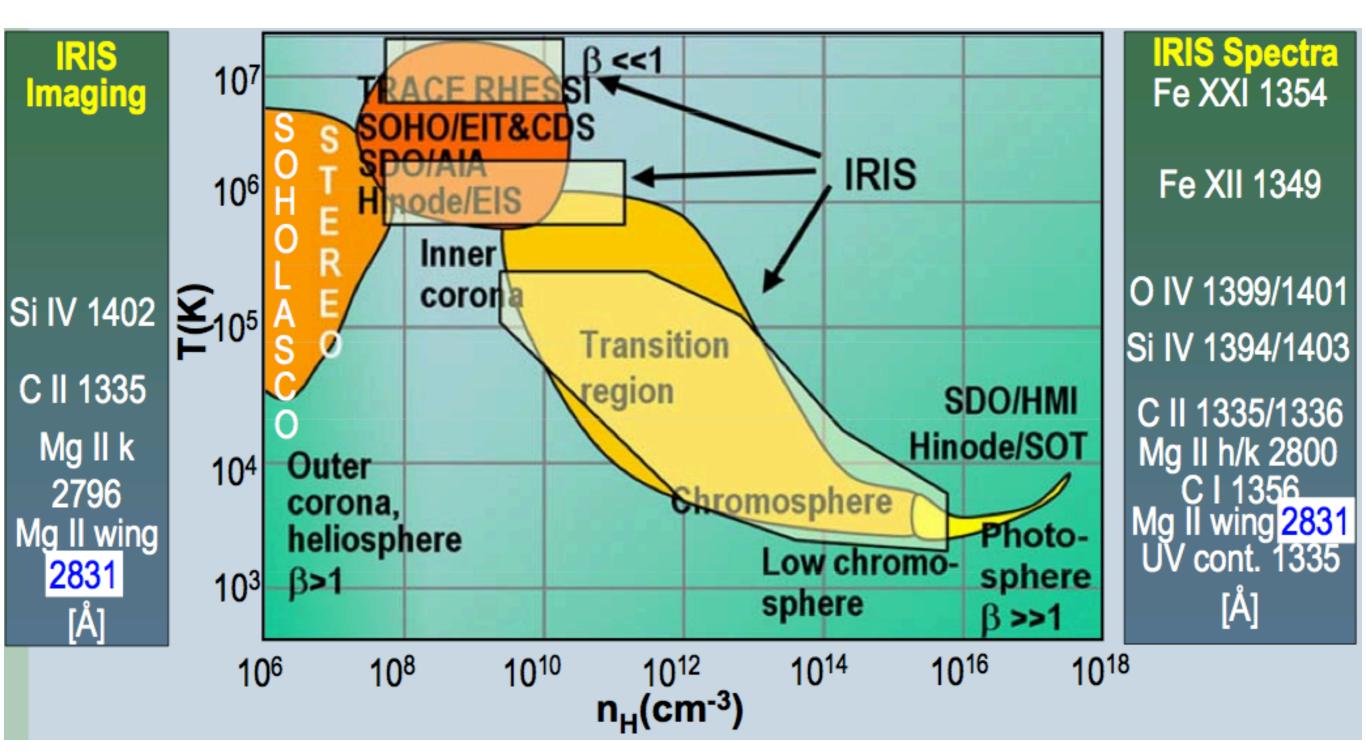
What is IRIS?

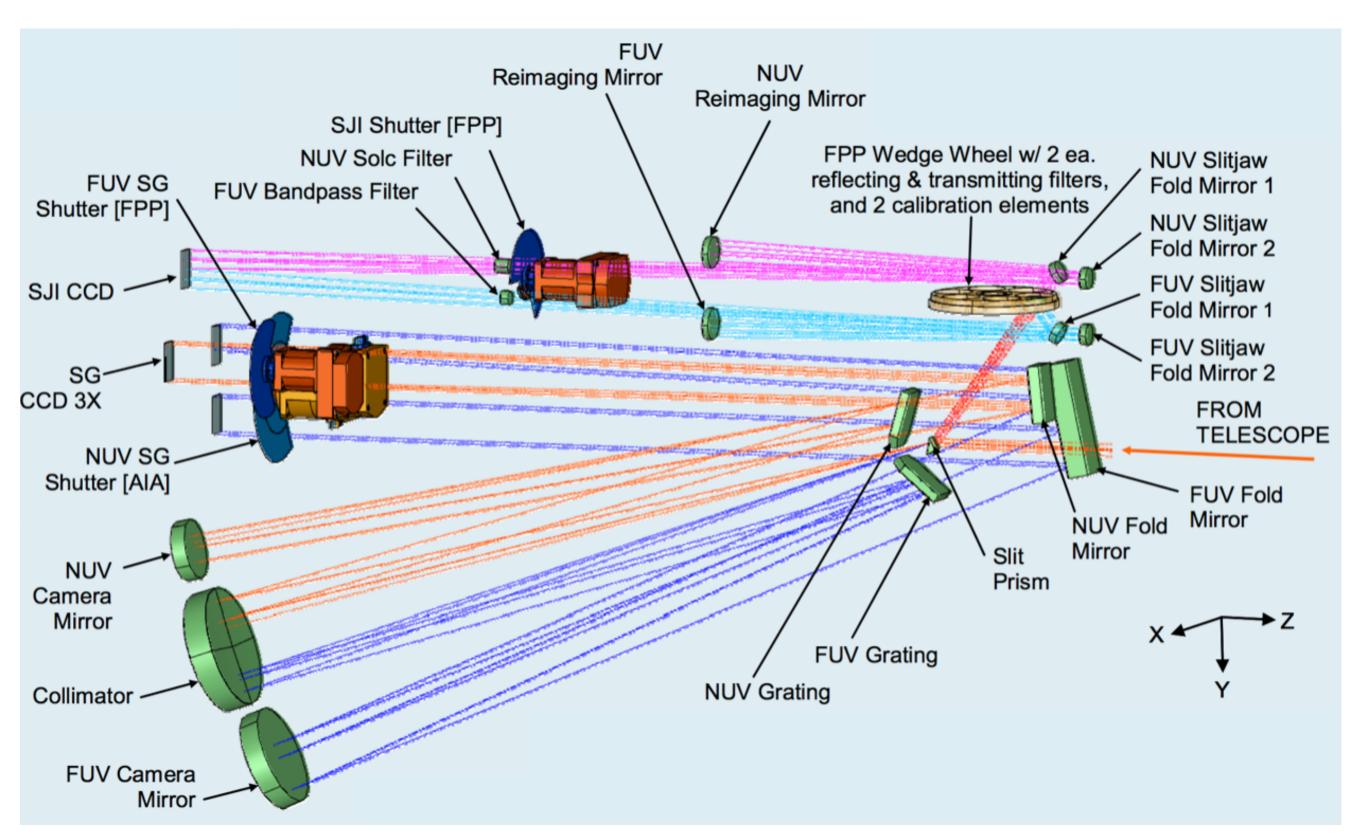
High resolution, far/near UV imaging spectrograph with slit-jaw imaging



Courtesy Bart De Pontieu

IRIS spectra and slit-jaw imaging cover the photosphere, chromosphere, transition region and corona - 4,500 to 10,000,000 K





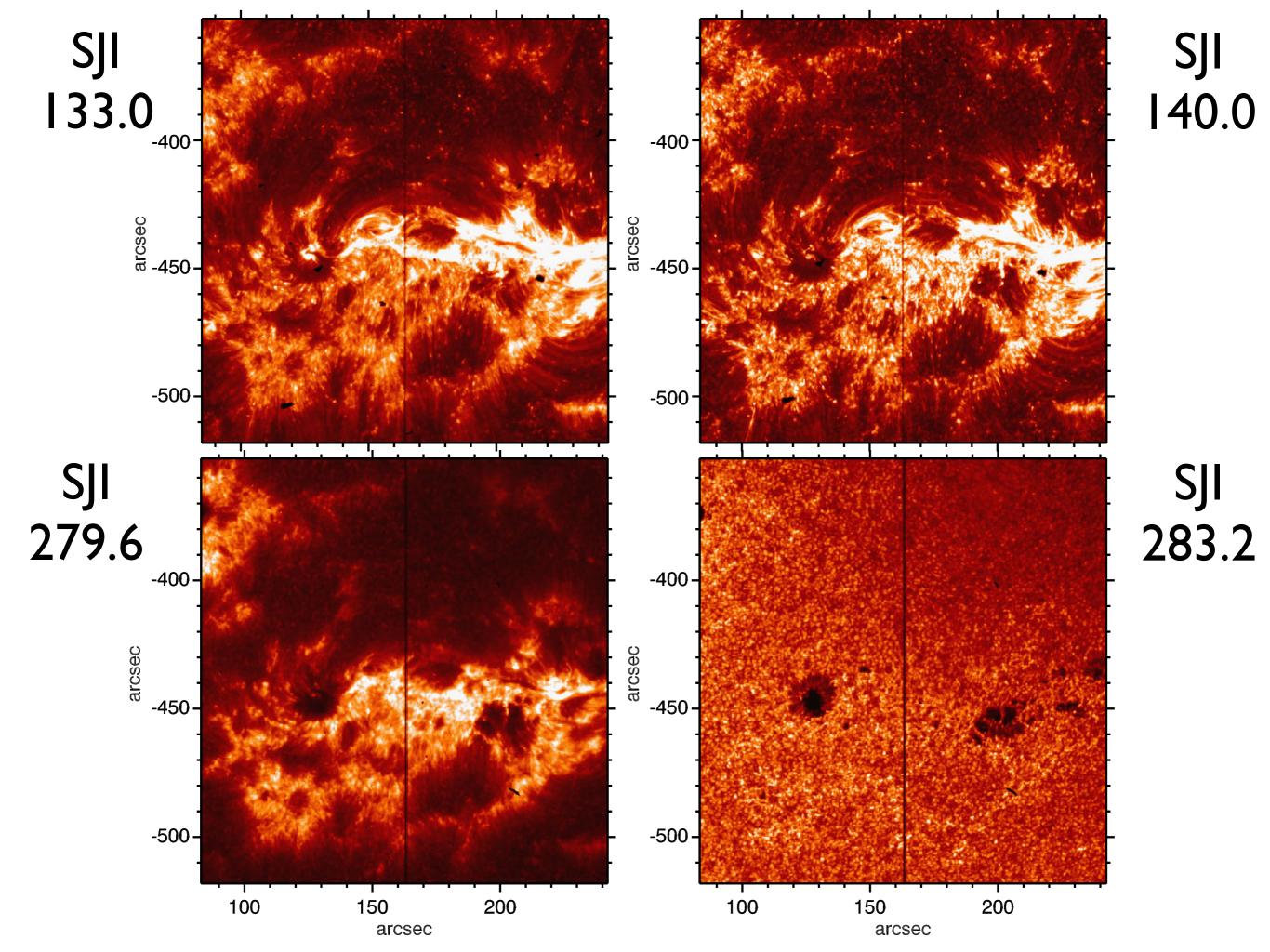
Schematic diagram of path taken by light in the FUV spectrograph (dark blue), NUV spectrograph (orange), FUV slit-jaw (light blue) and NUV slit-jaw (purple) path.

Table 2 IRIS spectrograph channels. Dispersion, Camera Electronics Box (CEB), and Effective Area (EA) vary for the three bandpasses.

Band	Wavelength [Å]	Disp. [mÅ pix ⁻¹]	FOV ["]	Pixel ["]	CEB	Shutter	EA [cm ²]	Temp. [log <i>T</i>]
FUV 1 FUV 2	1331.7 – 1358.4 1389.0 – 1407.0	12.98 12.72	175 175	0.1663 0.1663	1 1	FUV SG FUV SG	1.6 2.2	3.7 - 7.0 $3.7 - 5.2$
NUV	2782.7 – 2835.1	25.46	175	0.1664	2	NUV SG	0.2	3.7 - 4.2

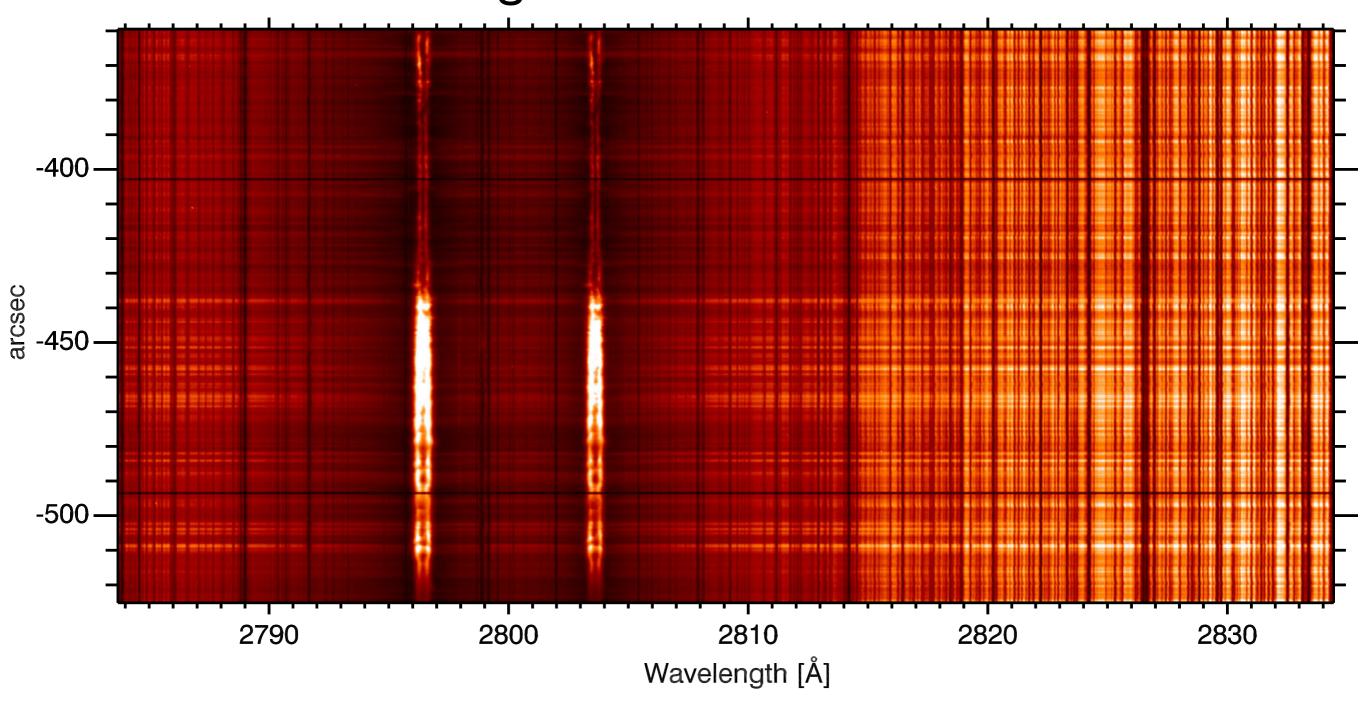
Table 3 IRIS slot channels. Filter-wheel positions can be either transmitting (T) or reflecting/mirrors (M).

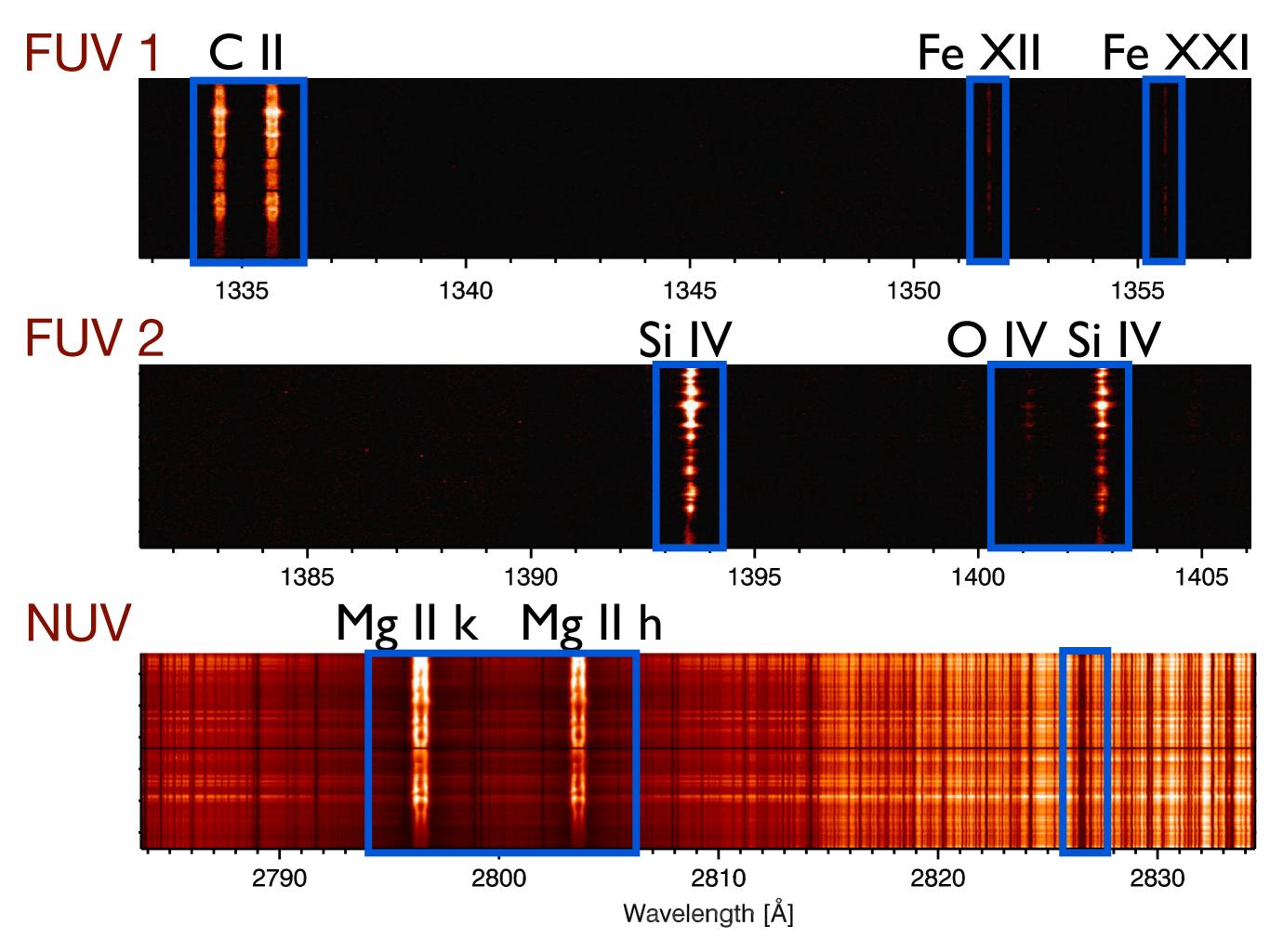
Band- pass	Filter wheel	Name	Center [Å]	Width [Å]	FOV ["×"]	Pix. ["]	EA [cm ²]	Temp. [log <i>T</i>]
Glass	1 T	5000	5000	broad	175 ²	0.1679	_	_
CII	31 M	1330	1340	55	175^2	0.1656	0.5	3.7 - 7.0
Mg II h/k	61 T	2796	2796	4	175^2	0.1679	0.005	3.7 - 4.2
Si IV	91 M	1400	1390	55	175^2	0.1656	0.6	3.7 - 5.2
Mg II wing	121 T	2832	2830	4	175^2	0.1679	0.004	3.7 - 3.8
Broad	151 M	1600W	1370	90	175^2	0.1656	_	



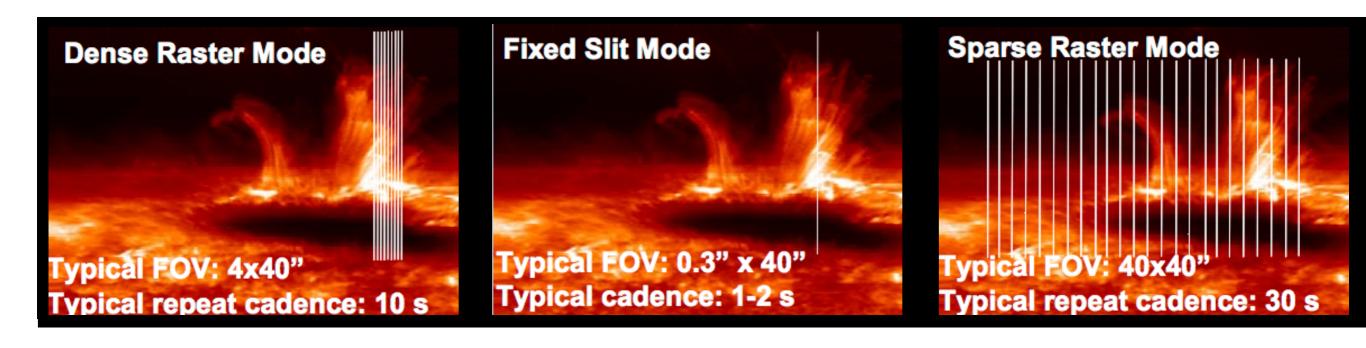
NUV spectra



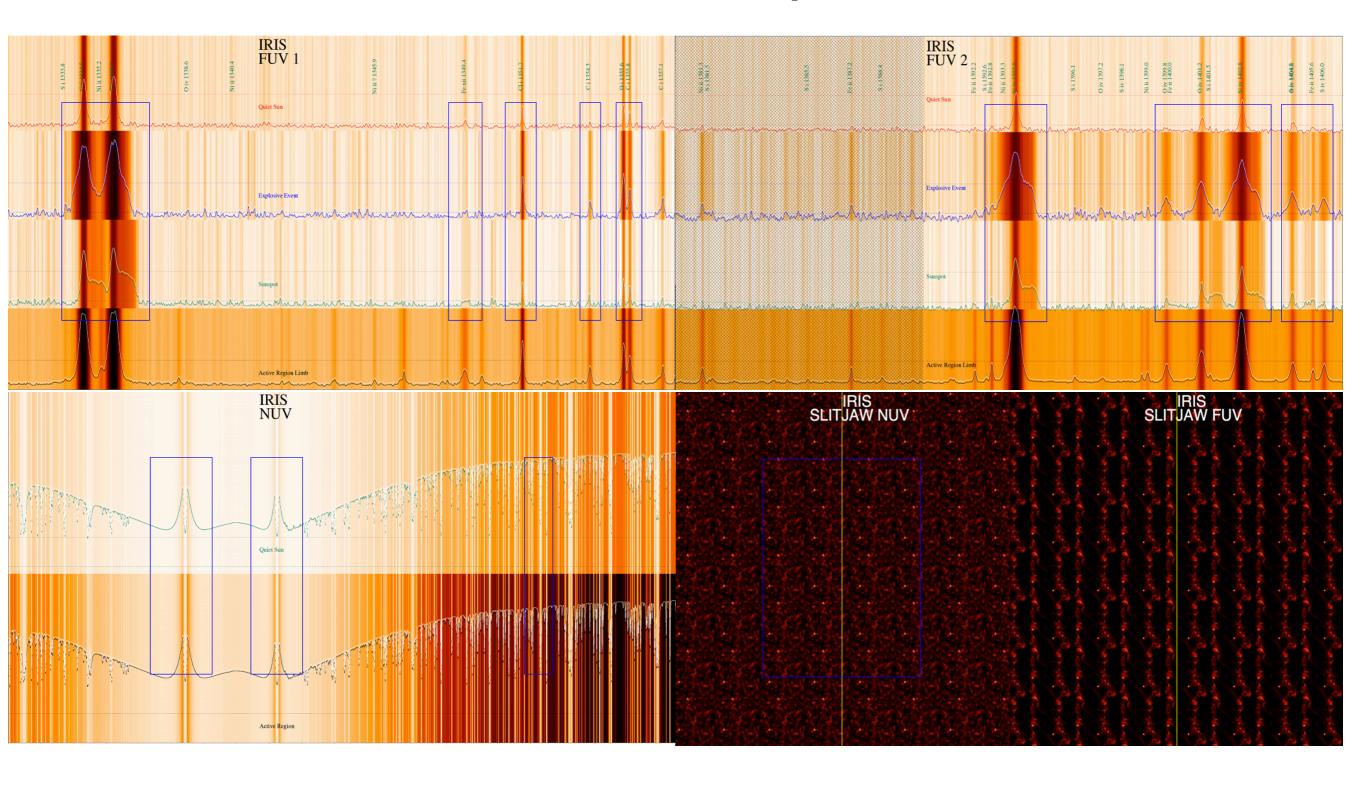




IRIS also performs sparse rasters to improve cadence (resulting in reduced data rate)



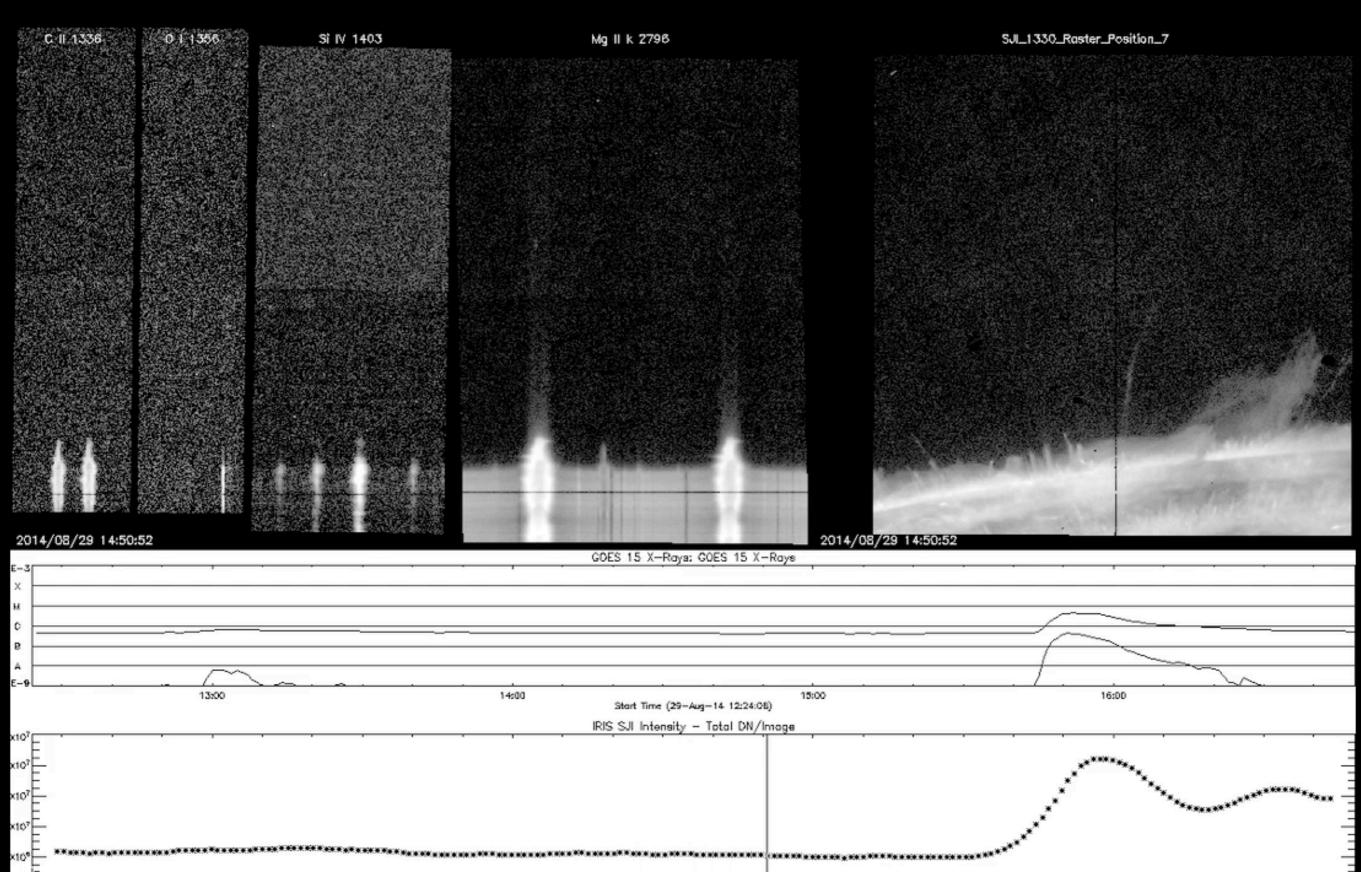
IRIS camera layout



CII OI Si IV 133.6 135.5 140.3



SJI 133 (C II + Fe XII)



Observing tables

OBS ID codifies the observing mode

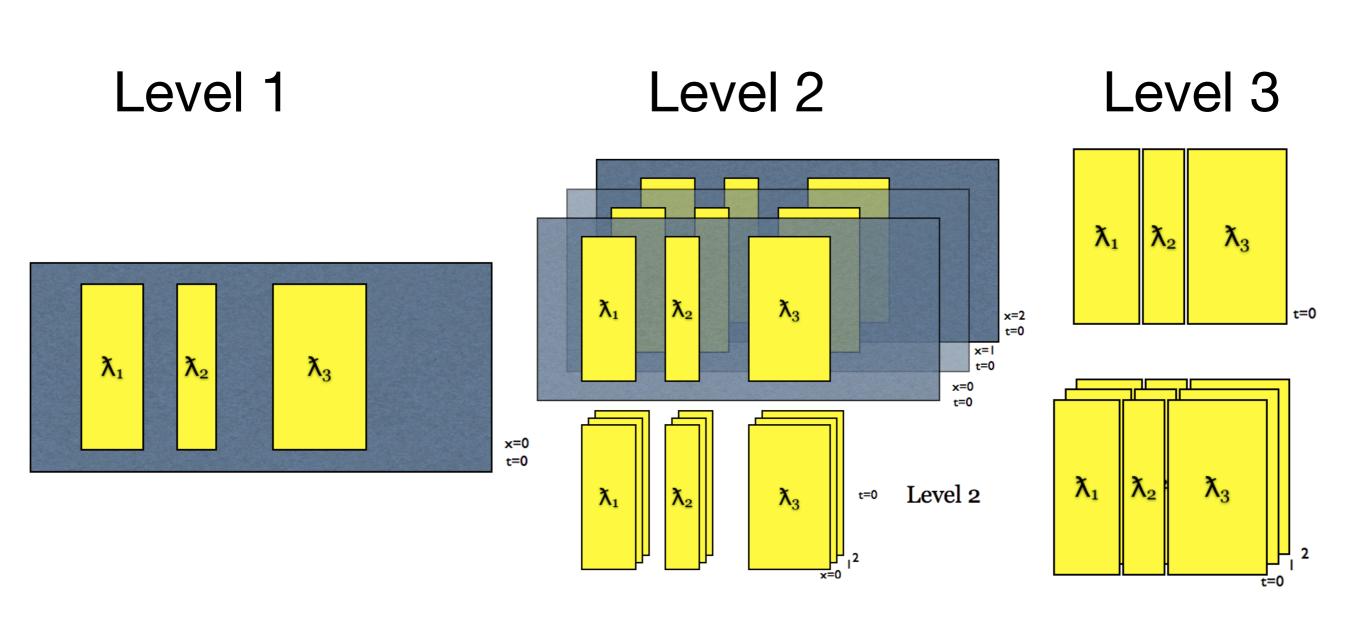
OBS ID parent	Description		
0-100	Basic raster type (sit-and-stare, rasters,)		
0-2,000	SJI choices		
0-12,000	Exposure times		
0-220,000	Summing modes (applied to FUV, NUV, SJI)		
0-750,000	FUV summing modes		
0-4,000,000	SJI cadence		
0-10,000,000	Compression choices		
0-180,000,000	Linelists		
3.8-4 billion	OBS table generation number		

See IRIS paper or ITN 31 for a detailed listing of the different modes.

IRIS data levels

Level	Processing	Notes		
TLM	Capture	Raw telemetry		
0	Depacketized	Raw images with basic keywords		
1	Reorient images to common axes: North up (0° roll), increasing wavelength to right	Lowest distributed level		
1.5	Dark current and offsets removed Flag bad pixels and pixels with spikes Flat-field correction Geometric and wavelength calibration	Transitory data product for level 2 production. Not distributed, for internal use only. Use iris_prep to go from level 1 to 1.5		
2	Recast as rasters and SJI time series	Standard science product. Scaled and stored as 16-bit integer.		
3	Recast as 4D cubes for NUV/FUV spectra.	CRISPEX format. May include transposed (sp) version. No SJI.		

IRIS data levels



INTERFACE REGION IMAGING SPECTROGRAPH

Home Mission Operations Data Analysis Modeling Documents Software Team Press Contact

IRIS mission/instrument paper

Operations/Planning

ITN 1 - IRIS Operations Overview

ITN 2 - Manual for Table Creator

ITN 3 - Manual for Timeline Tool

ITN 4 - Manual for Synthetic Observations Tool

ITN 5 - Operations Under Roll Conditions

ITN 6 - AEC Operations

ITN 7 - Compression Approach

ITN 8 - Checklist for IRIS planner

ITN 9 - Periodic Calibration Activities

Data Flow

ITN 10 - General Approach to Data Flow and Archiving

ITN 11 - Definition of Data Levels

ITN 12 - Definition of Keywords

ITN 13 - VSO and IRIS

Calibration

ITN 14 - Dark Current/Offset

ITN 15 - Despiking

ITN 16 - Flat-field

ITN 19 - Geometric Calibration

ITN 20 - Wavelength Calibration

ITN 21 - Recasting into Level 2/3 Data

ITN 22 - Co-alignment, Plate Scale Analysis

ITN 23 - MTF/PSF Determination

ITN 24 - Stellar Calibration

ITN 25 - Gain Determination

Data Analysis

ITN 26 - User Guide To Data Analysis

ITN 27 - Quicklook Tools Manual

ITN 28 - IRIS IDL Data Structure

ITN 29 - Deconvolution Approach

ITN 30 - 60 Day Observing Plan

ITN 31 - IRIS science planning: tables, linelists, targets

SolarSoft Tree and UVSP Database

Numerical Modeling

ITN 33 - General Overview of Numerical Simulations

ITN 34 - Numerical Simulations Quicklook Tools

ITN 35 - Numerical Simulations Synthetic Observables

ITN 36 - RH 1.5 D Manual

ITN 37 - How to Derive Physical Information from Mg II h/k

IRIS Technical Notes List (ITN)

Talks & Posters

Invited Talks at the SDO-4/IRIS/Hinode Workshop - March 2012 IRIS Talk Hinode 5 Meeting Keynote (430 MB) PDF (36 MB)

IRIS Poster

FUV Camera View

NUV Camera View

Concept Study Report

Executive Summary

Science Goals

Instrument Description

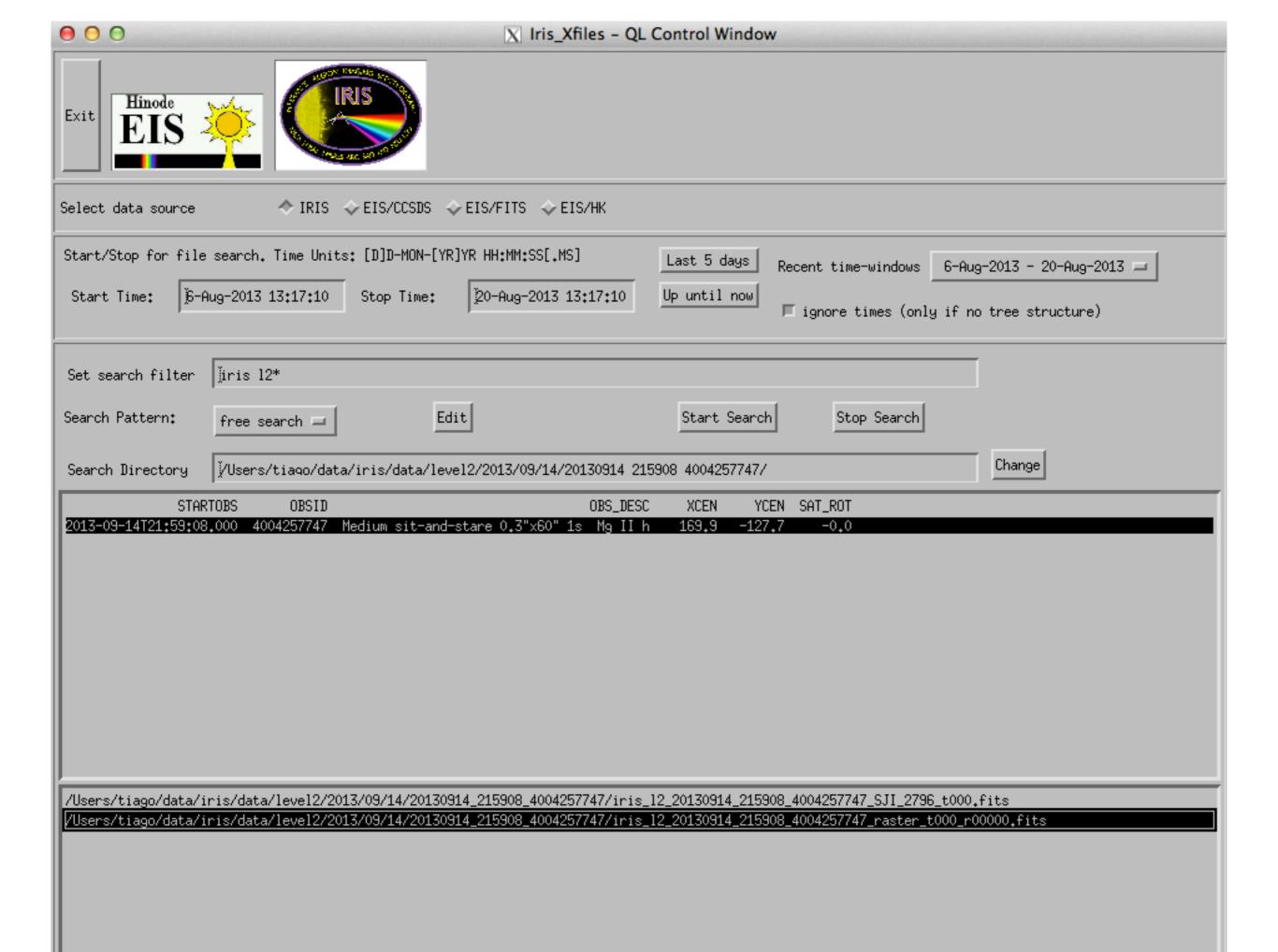
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Searching, downloading, browsing data

- IRIS search webpage http://iris.lmsal.com/search/
- Hinode SDC Europe http://sdc.uio.no/search/API
- SolarSoft IDL
- IRIS today: http://iris.lmsal.com/iristoday/
- HEK recent observations: <u>http://www.lmsal.com/hek/hcr?cmd=view-recent-</u> events&instrument=iris

Live demo: searching and downloading



Live demo: IRIS xfiles

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IDL object interface for level 2 data

```
IDL> f = 'iris_l2_20131010_100202_3820259146_raster_t000_r000000.fits'
IDL> d = iris_obj(f)
```

```
IDL> d->show_lines
Spectral regions(windows)
0  1335.71    C II 1336
1  1349.43    Fe XII 1349
2  1355.60    O I 1356
3  1393.78    Si IV 1394
4  1402.77    Si IV 1403
5  2832.76  2832
6  2814.50  2814
7  2796.20    Mg II k 2796
```

Read IRIS L2

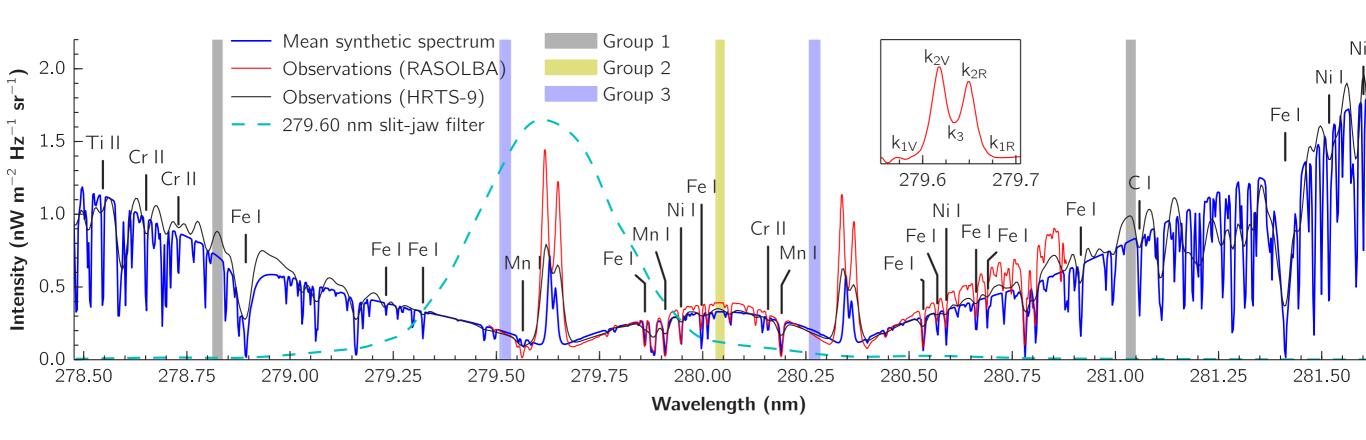
```
IDL> sjifile = 'iris_l2_20131010_100202_3820259146_SJI_2796_t000.fits'
IDL> read_iris_l2, sjifile, header, data
(...)
IDL> help, header, data
HEADER STRUCT = -> <Anonymous> Array[100]
DATA FLOAT = Array[1860, 1092, 100]
```

Live demo: Read IRIS data in IDL

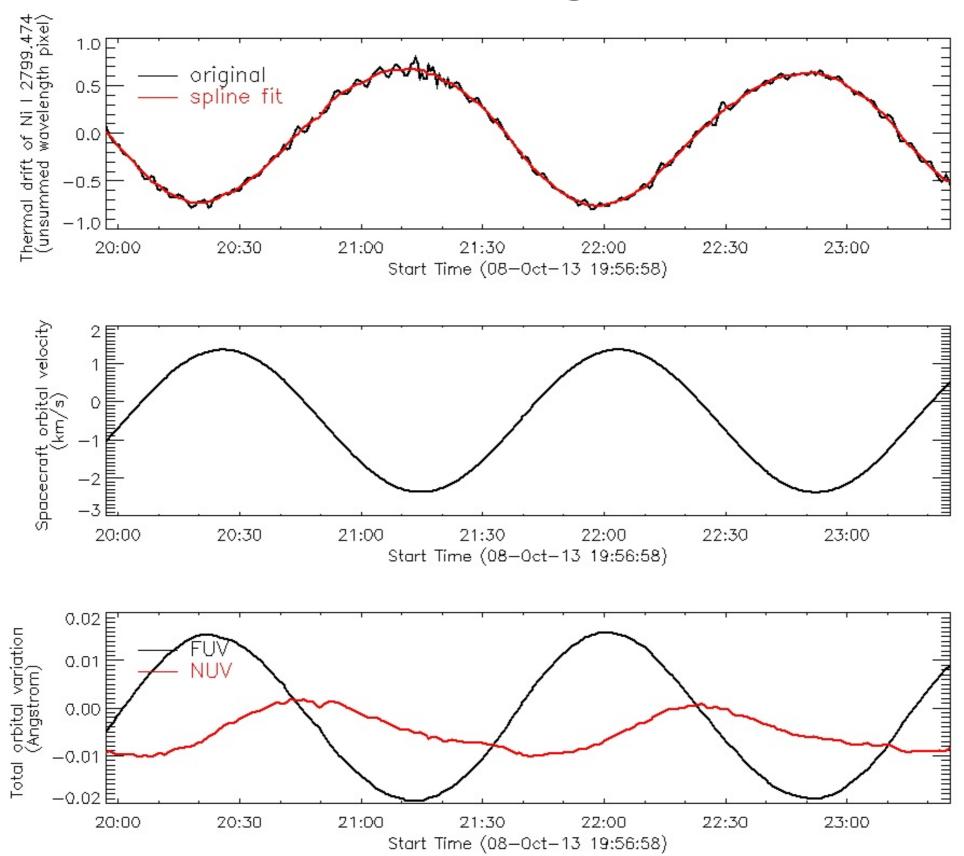
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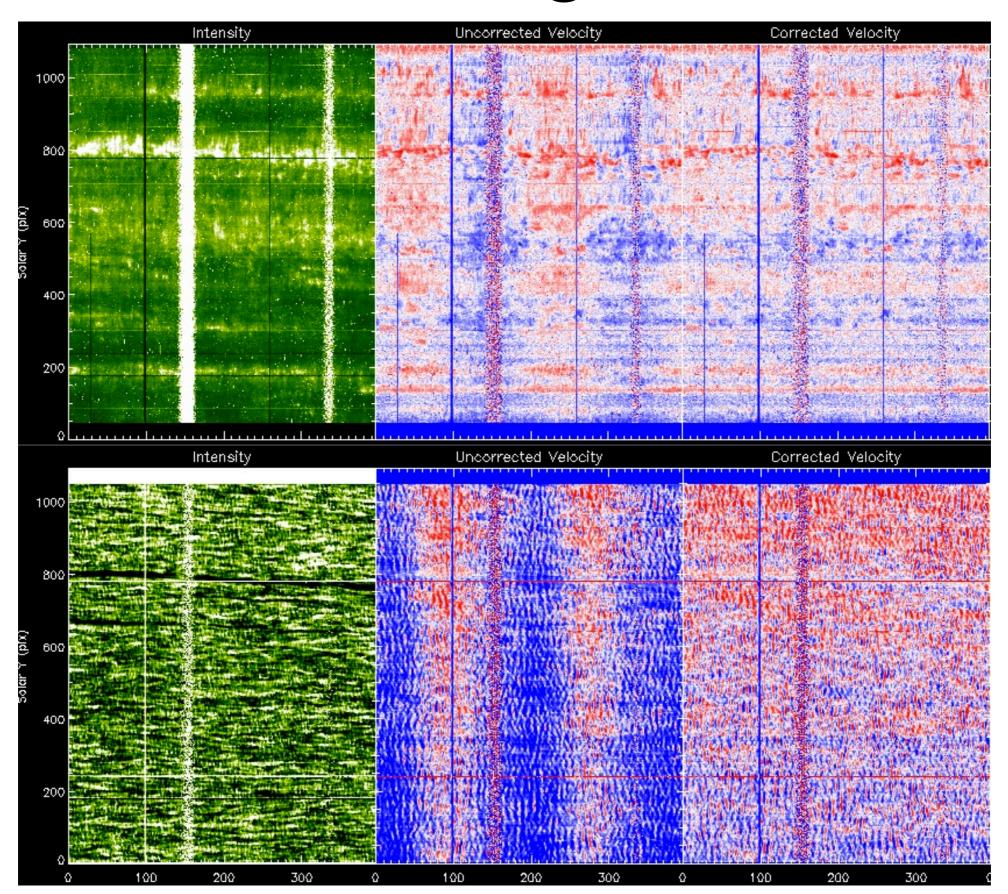
Precise wavelength calibration

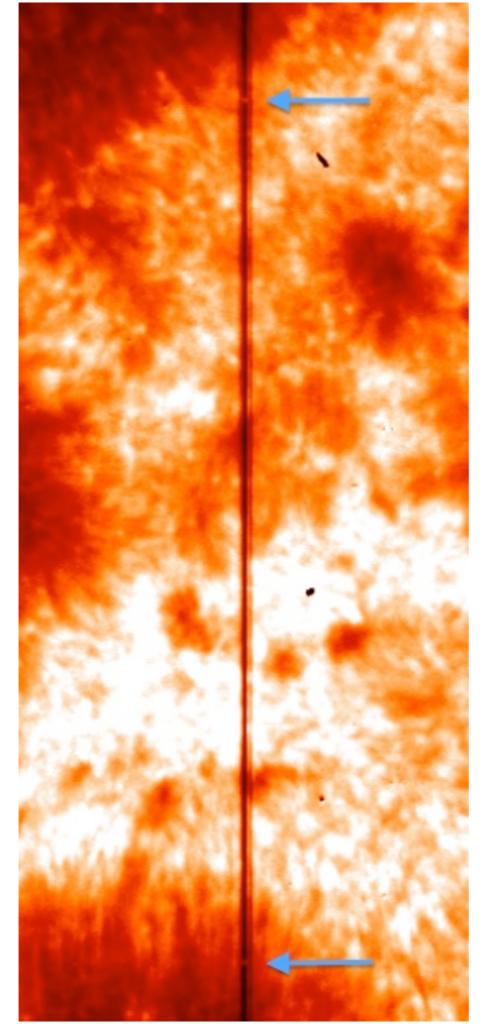


Precise wavelength calibration



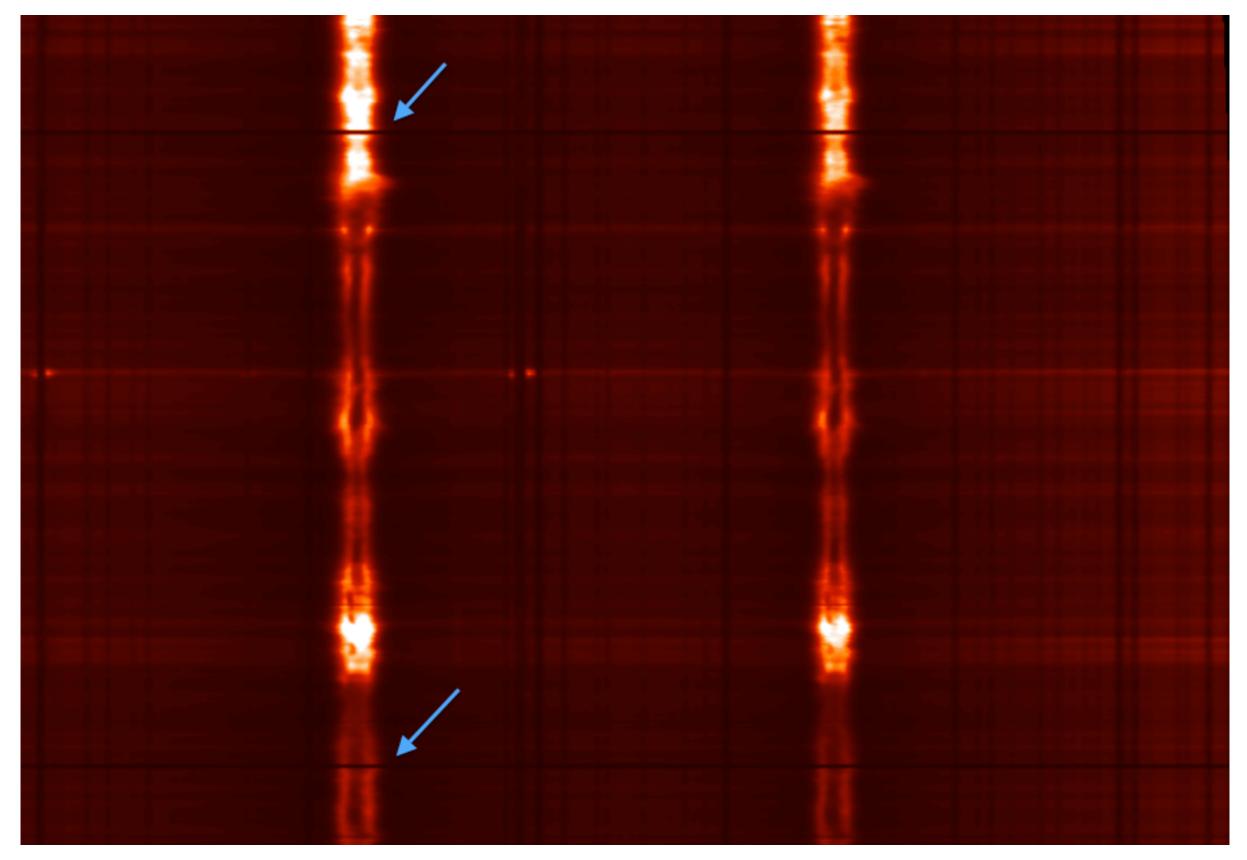
Precise wavelength calibration





Co-alignment between SJIs

Co-alignment between spectra



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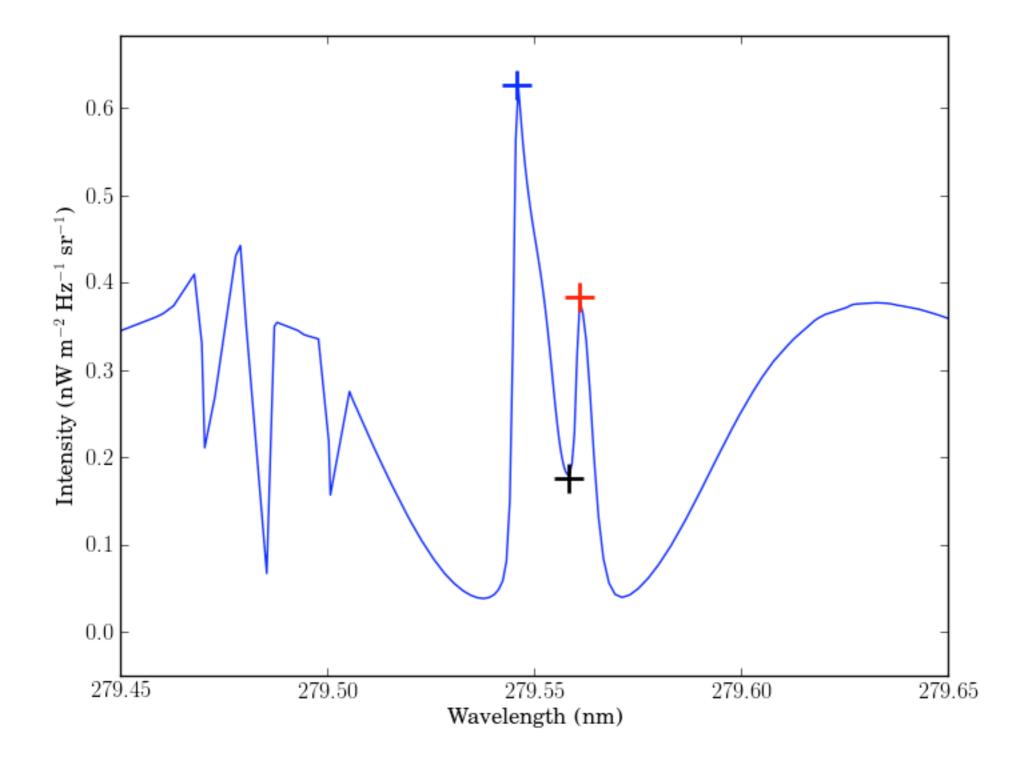
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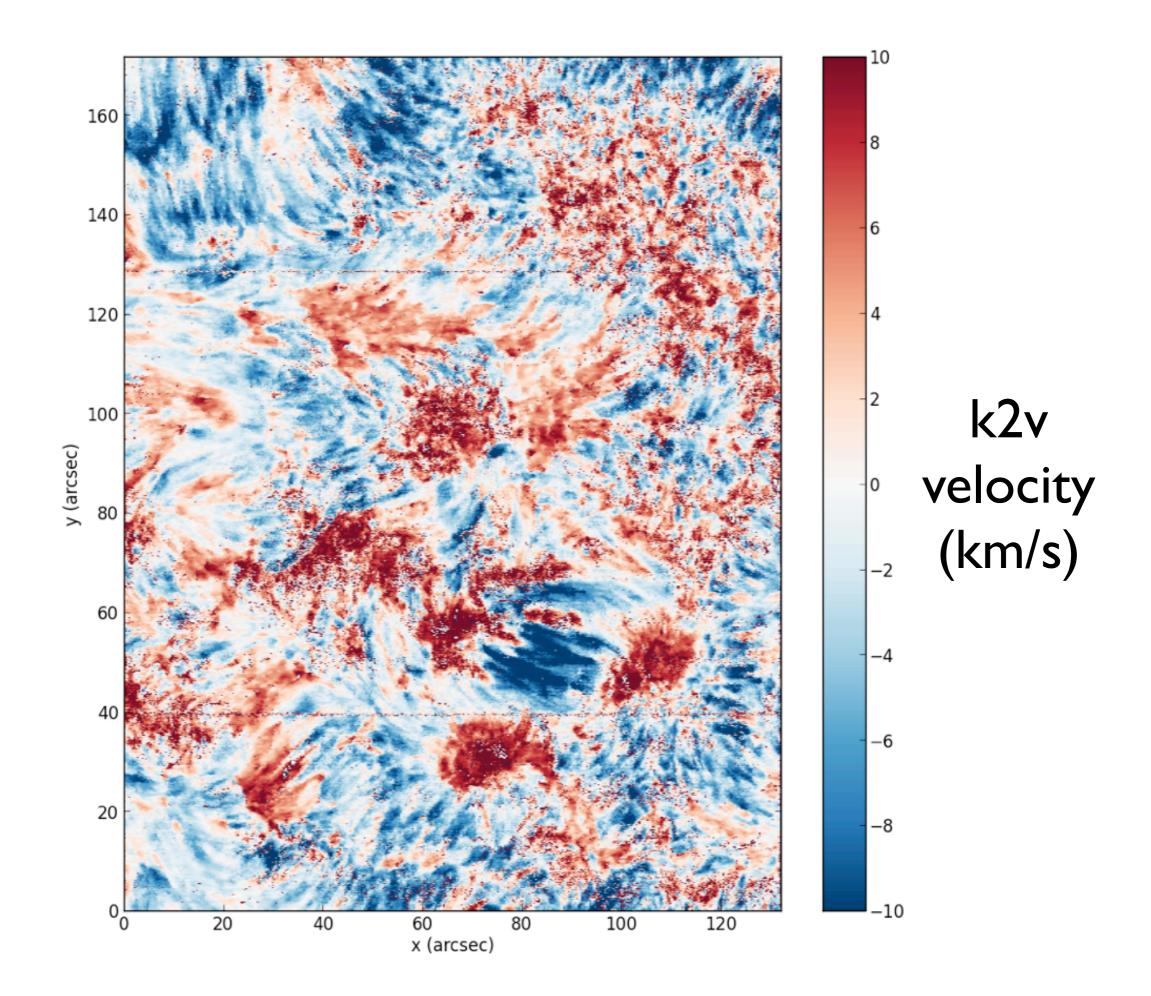
Exercise questionsHands-on tutorials

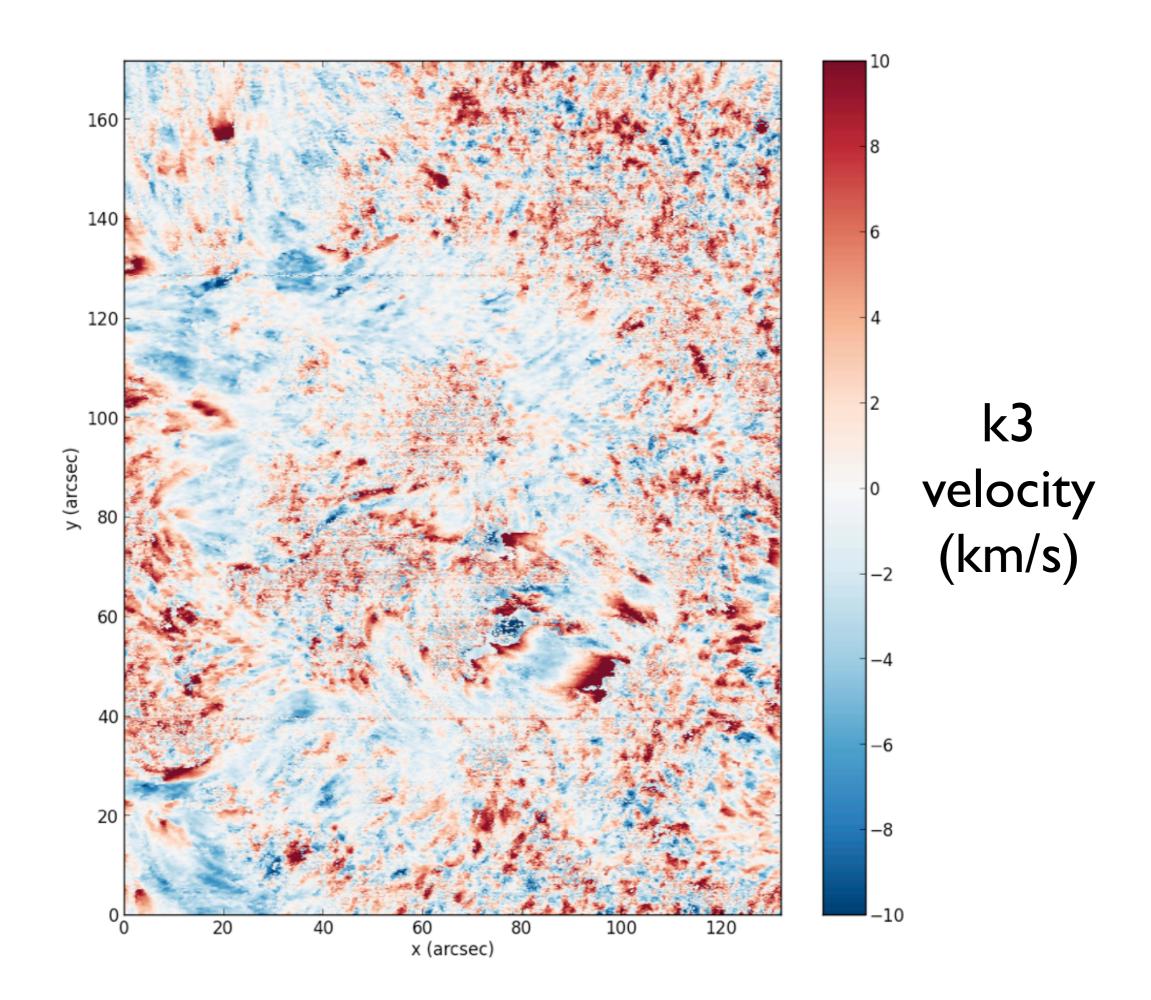
Algorithm to automatically measure Mg II line properties available on solarsoft:

iris_get_mg_features_lev2, file, mg_id, vr, lc, rp, bp

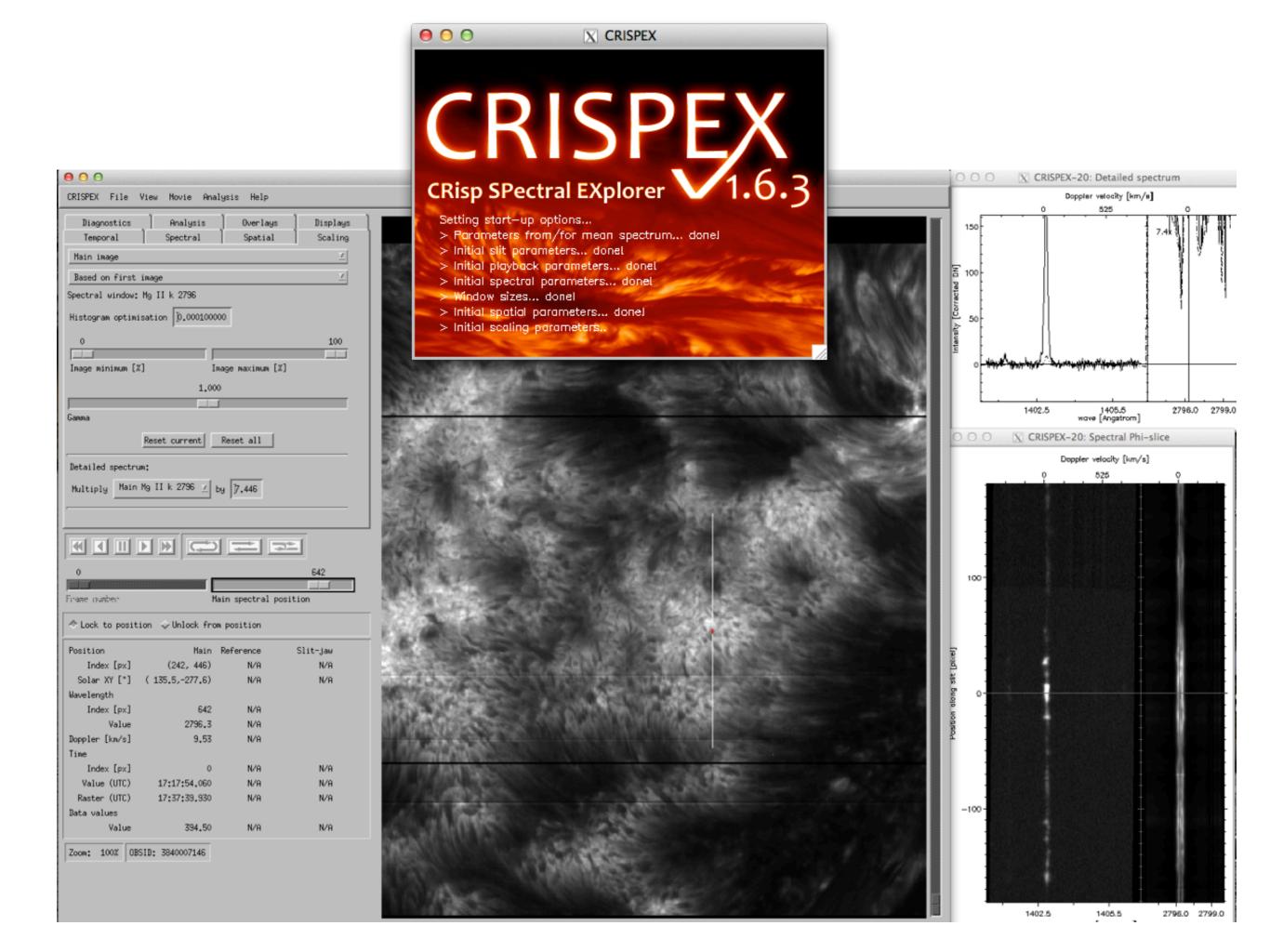
Based on the algorithm described in Pereira et al. (2013)







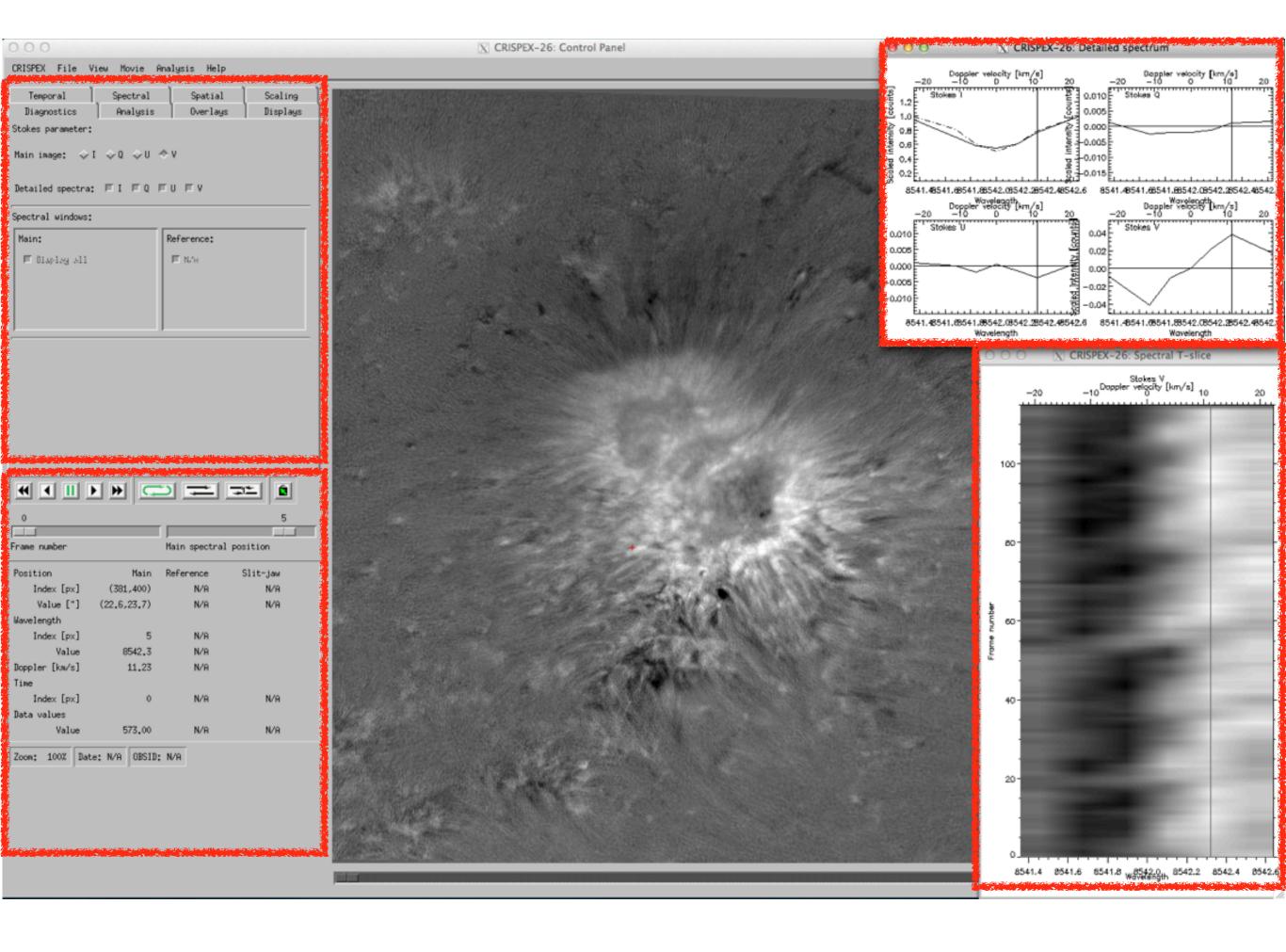
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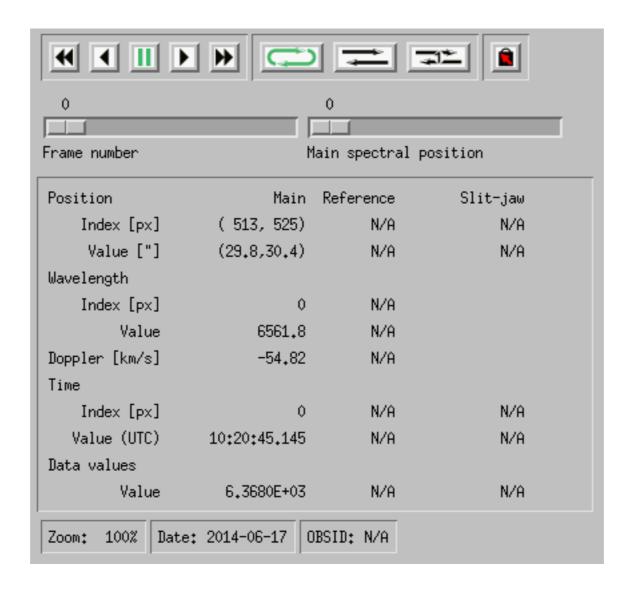
CRISPEX file formats

- "La Palma cubes"
 - ★ Simple cubes of (nx, ny, nwave * nt * nstokes)
 - ★ Combined with "spectfile"
- IRIS level 3 fits files
 - ★ FITS file with main image (nx, ny, nwave, nt)
 - ★ FITS keywords used for coordinates, time
 - ★ Extensions with wavelength and time values
 - ★ Not limited to IRIS data; to be further standardised

Two types of files: (same data) 'im' (nx, ny, nw, nt) and 'sp' (nw, nt, nx, ny)

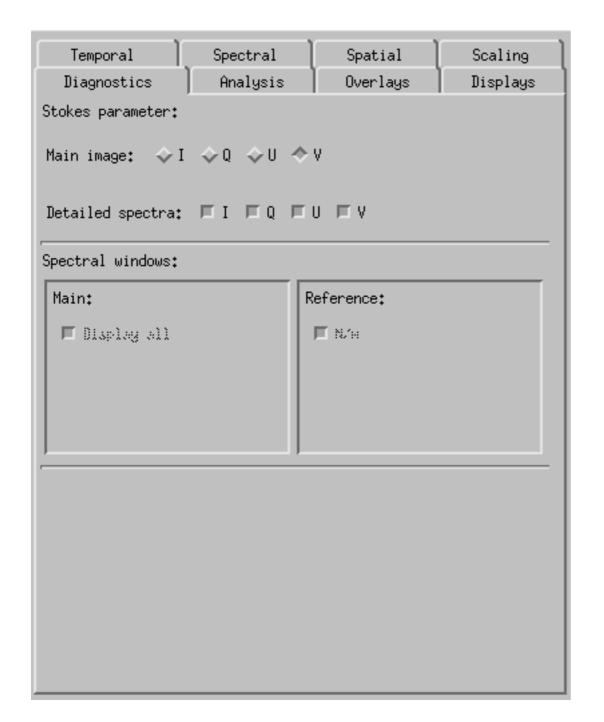


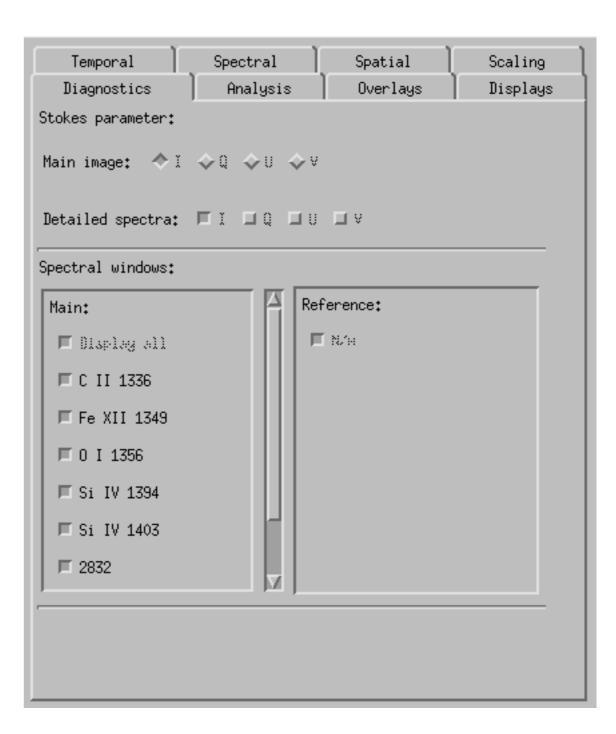
Bottom control panel



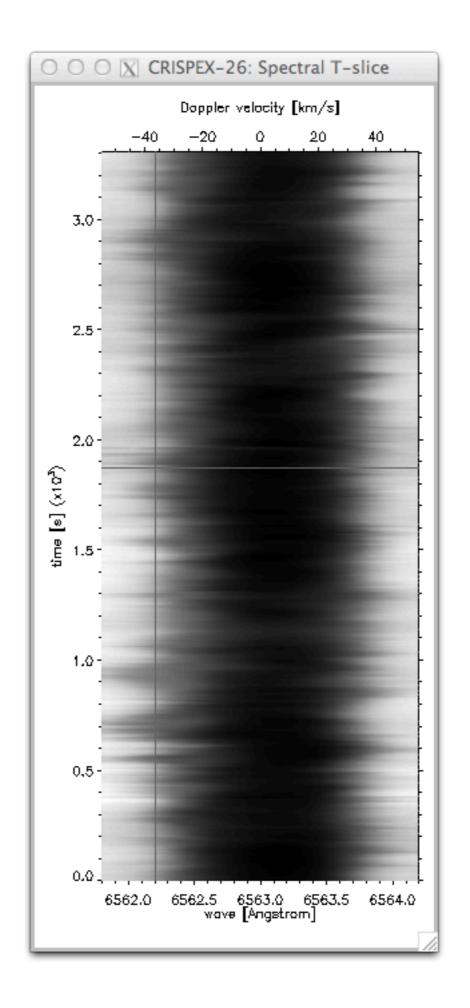
Tabs

Diagnostics Analysis Overlays Displays Temporal Spectral Spatial Scaling				
Lower index: Upper index: \$598 Reset				
Update spectral windows				
10				
Animation speed [frame/s]				
1				
Frame increment				
⊒ Blink between waln and reference iwage				
Master time: Main				
0				
Easter timing offset [raster position]				





Diagnostics Temporal	Analysis Spectral	Overlays Spatial	Displays Scaling	
Main image				
Based on first image				
Spectral window: Halpha SST				
Histogram optimisation 0.000100000				
0			100	
Image minimum [%] Image maximum [%]				
1,000				
Gamma				
Reset current Reset all				
Detailed spectrum:				
Multiply Main Halpha SST ∠ by \$.779				



Live demo: CRISPEX

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http://pollev.com/iris7



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